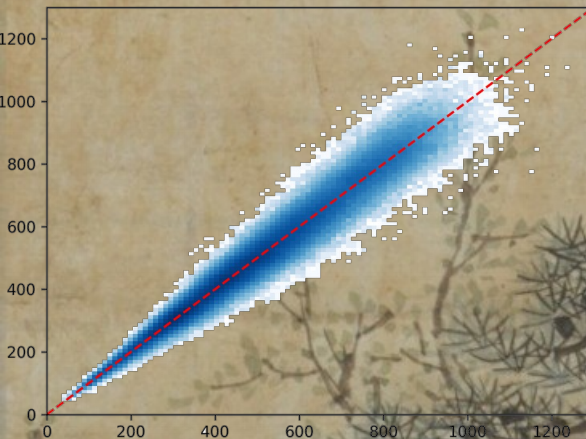
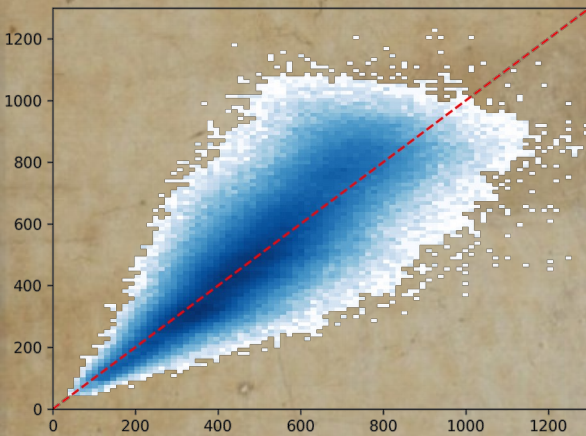


# Peculiar velocities derived from the Fundamental Plane

by

Christoph Saulder



KIAS

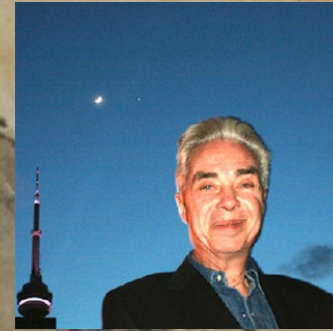
KOREA  
INSTITUTE FOR  
ADVANCED  
STUDY



➤ Christoph Saulder (KIAS)



➤ Ian Steer (NED)



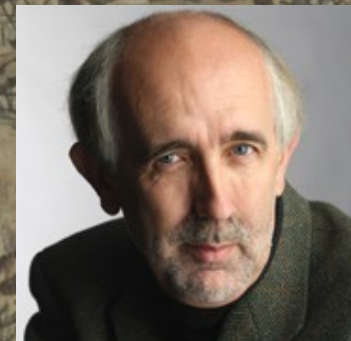
➤ Owain Snaith (KIAS)



➤ Changbom Park (KIAS)



➤ Barry F. Madore (NED)

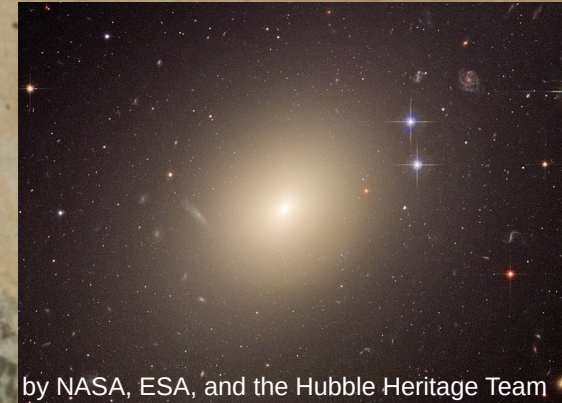




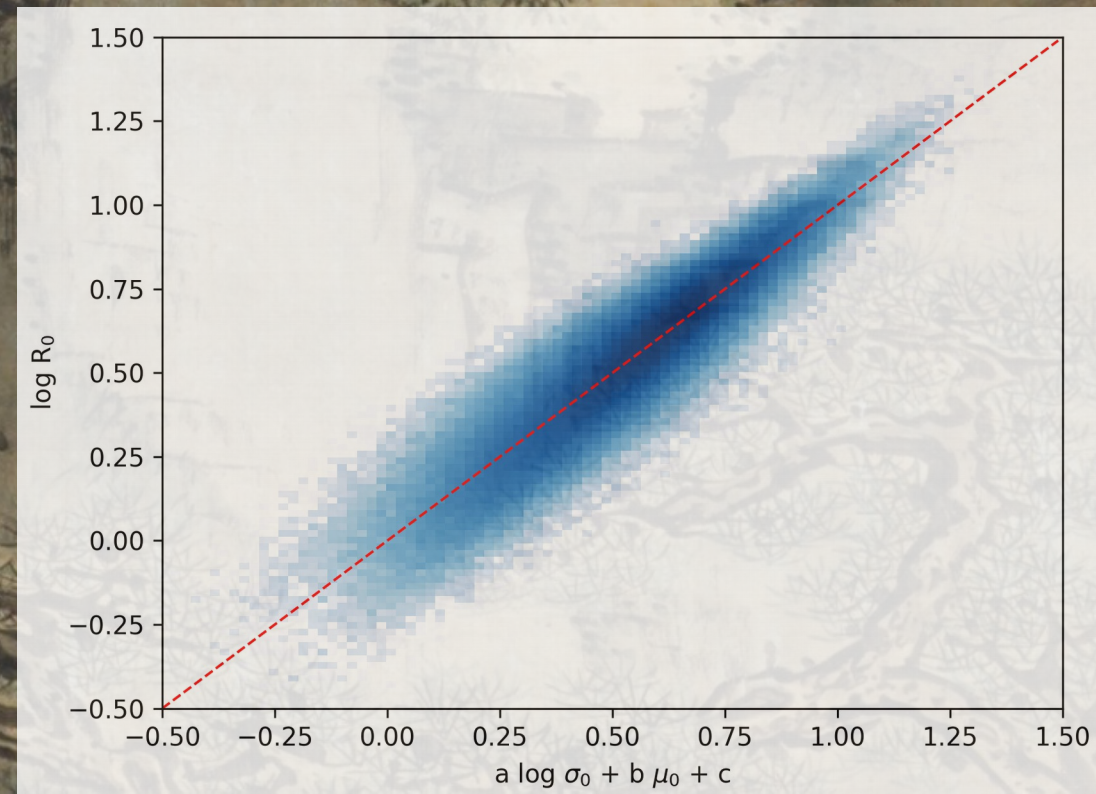
# The traditional fundamental plane

- Empirical distance indicator for early-type galaxies:

$$\log(R_0) = a \log(\sigma_0) + b \mu_0 + c$$



- ~20% accuracy
- Assuming:
  - luminosity-independent M/L ratios for ETG
  - virial equilibrium
- $a=2$ ,  $b=0.4$ , but in reality  $a \sim 1$ ,  $b \sim 0.3$





# Central velocity dispersion data is rare

- Central velocity dispersions are obtained by fibre spectroscopy
- Large samples of good quality data are rare:
  - 6dFGS follow-up (6dFGSv)
  - SDSS/BOSS
  - Taipan Survey (ongoing)
- SDSS/BOSS provides the by far largest sample (and also the photometric data)



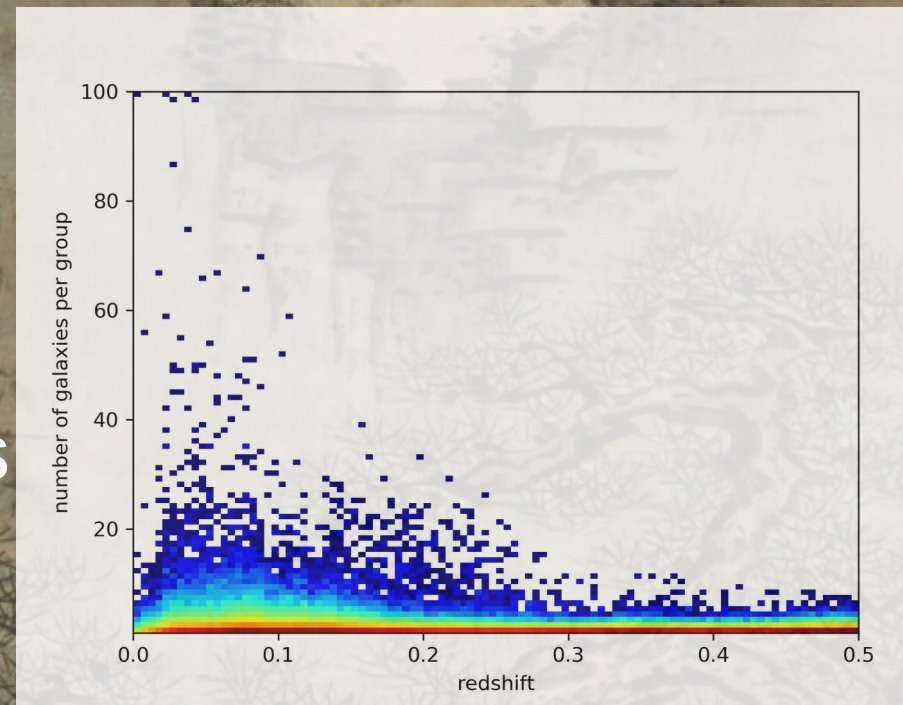
# The largest sample of fundamental plane data

- Notable previous (large) calibrations:
  - Bernardi+ 2003: ~8 000 ETG in SDSS
  - Hyde&Bernardi 2009: ~46 500 ETG in SDSS
  - La Barbera+ 2010: ~4 500 ETG in SDSS
  - Magoulas+ 2012 & Campbell +2014: ~9 000 ETG in 6dFGS
  - Saulder+ 2013: ~93 000 ETG in SDSS
  - Saulder+ 2015&2016: ~119 000 ETG in SDSS
- NOW: ~280 000 early-type galaxies identified in SDSS DR 14 (including BOSS)



# Group catalogue

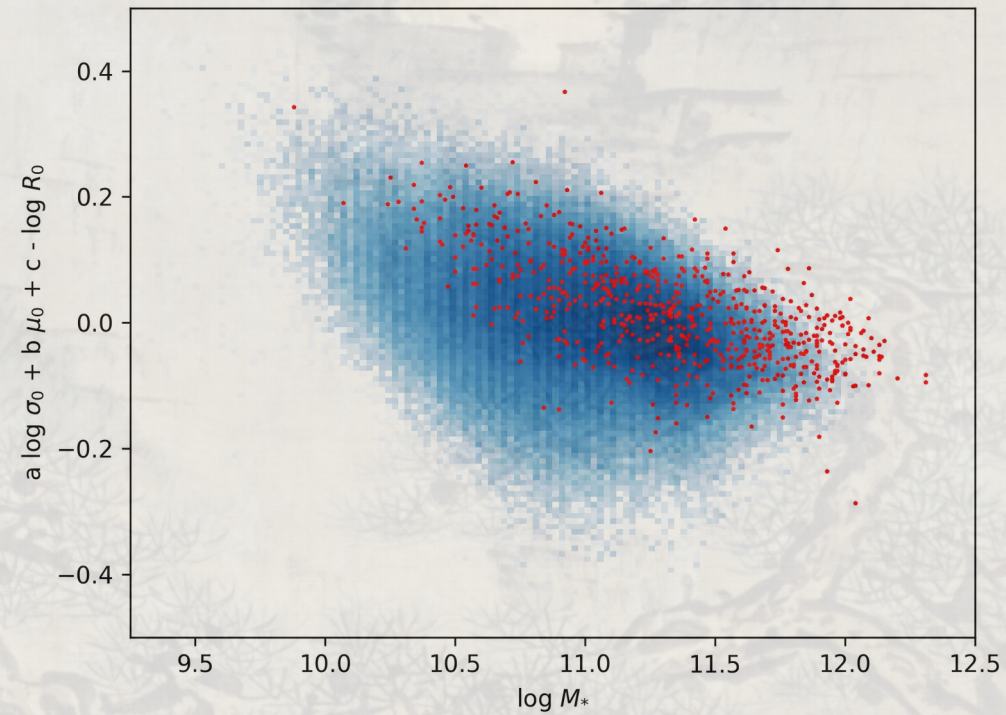
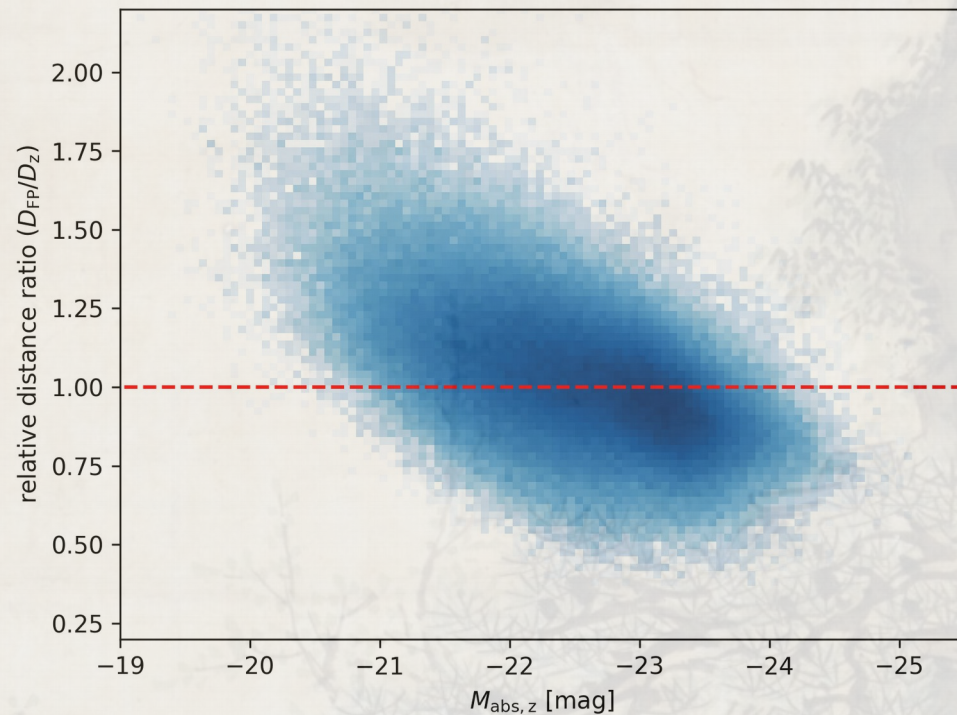
- Combined SDSS/BOSS Friends-of-Friends group catalogue up to  $z=0.5$
- 2MRS included to compensate for the saturation BIAS of SDSS spectroscopy
- Linking lengths calibrated used mock catalogues based on the WMAP7 re-run of the Millennium simulation
- Collapsing the “finger of God” effect for clusters
- Improving FP distance estimates to groups hosting several ETG
- Comparing FP-distances to Tully-Fisher relation distances





# Problematic biases

- Strong dependence on absolute magnitude → stellar mass dependence (very clear with MaNGA)
- Assumption that M/L is luminosity independent is wrong → more than just a tilt
- Serious issue for magnitude limited surveys





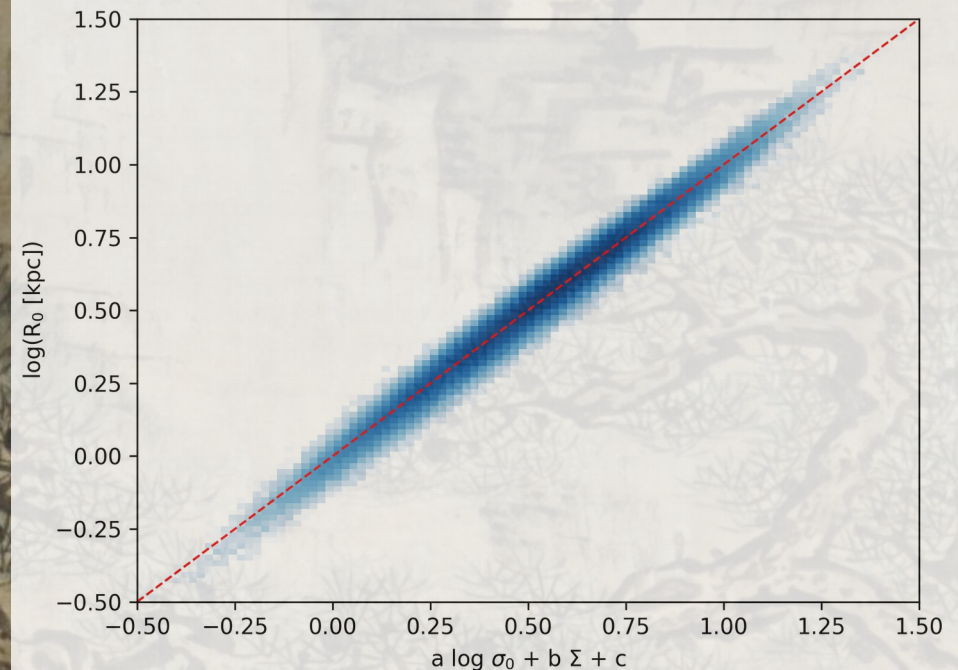
# The stellar mass fundamental plane

- Discussed in Hyde&Bernardi+2009, but not used to its full potential (as a distance estimator):

$$\log(R_0) = a \log(\sigma_0) + b \log(\Sigma) + c$$

$$\text{with } \log(\Sigma) = \log(M_*) + m_{\text{abs}} - \mu_0 = \log(M_*/L) - \mu_0 + k$$

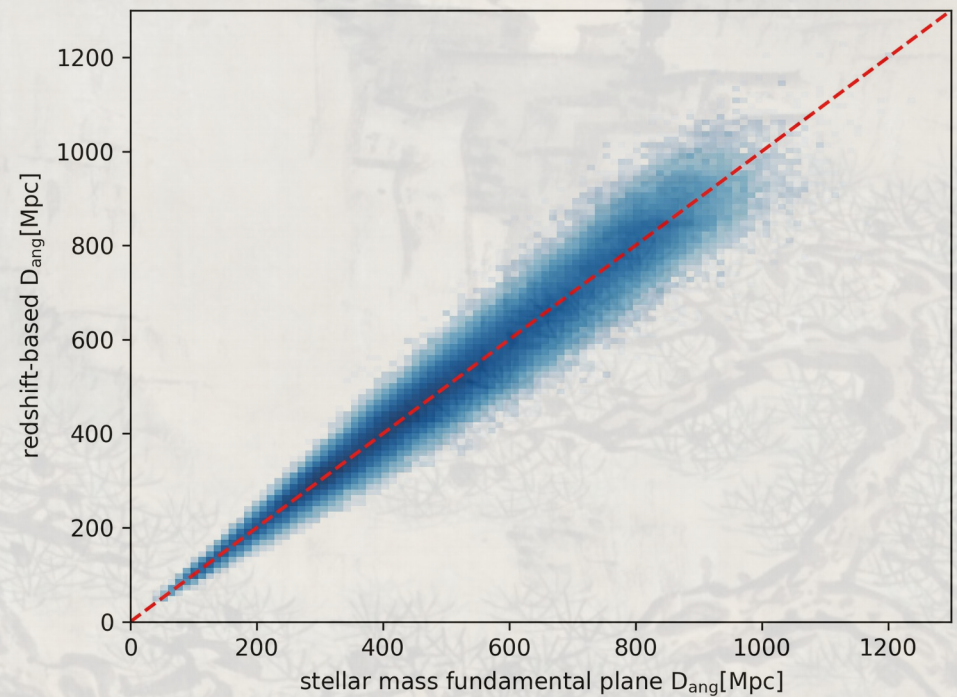
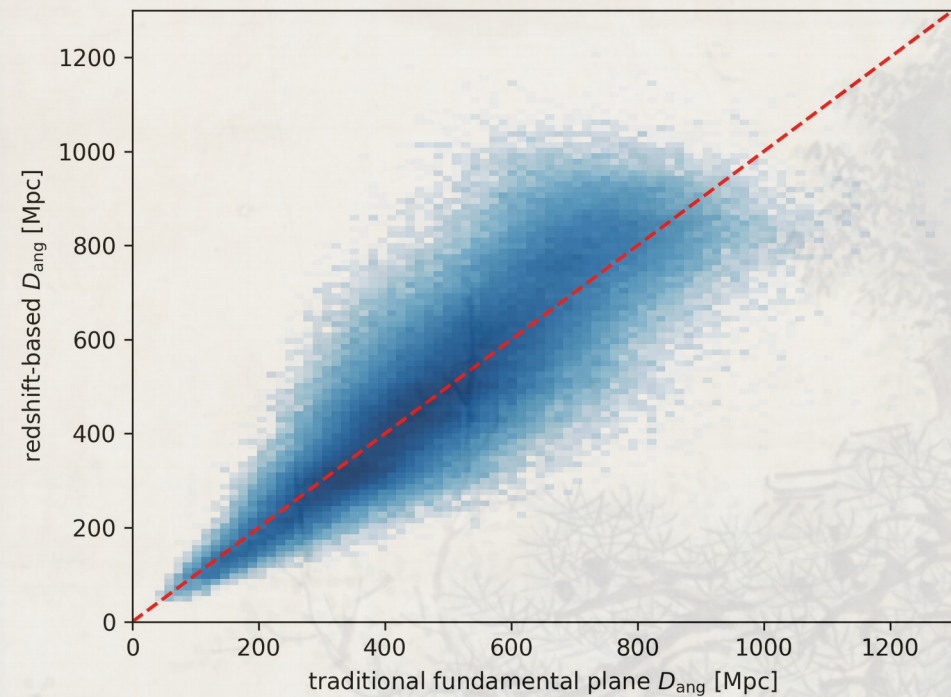
- Has a hidden redshift dependence ( $M_*$  and  $m_{\text{abs}}$ )
- Tempered by using their combination:  $M_*/L$
- Trading biases and gains





# Improvement in distance measurements

- Group catalogue improves average accuracy by 0.5%
- Traditional FP: 19.8% accuracy on average
- Stellar mass FP: 7.7% (stat.) + 0.8% (sys.)





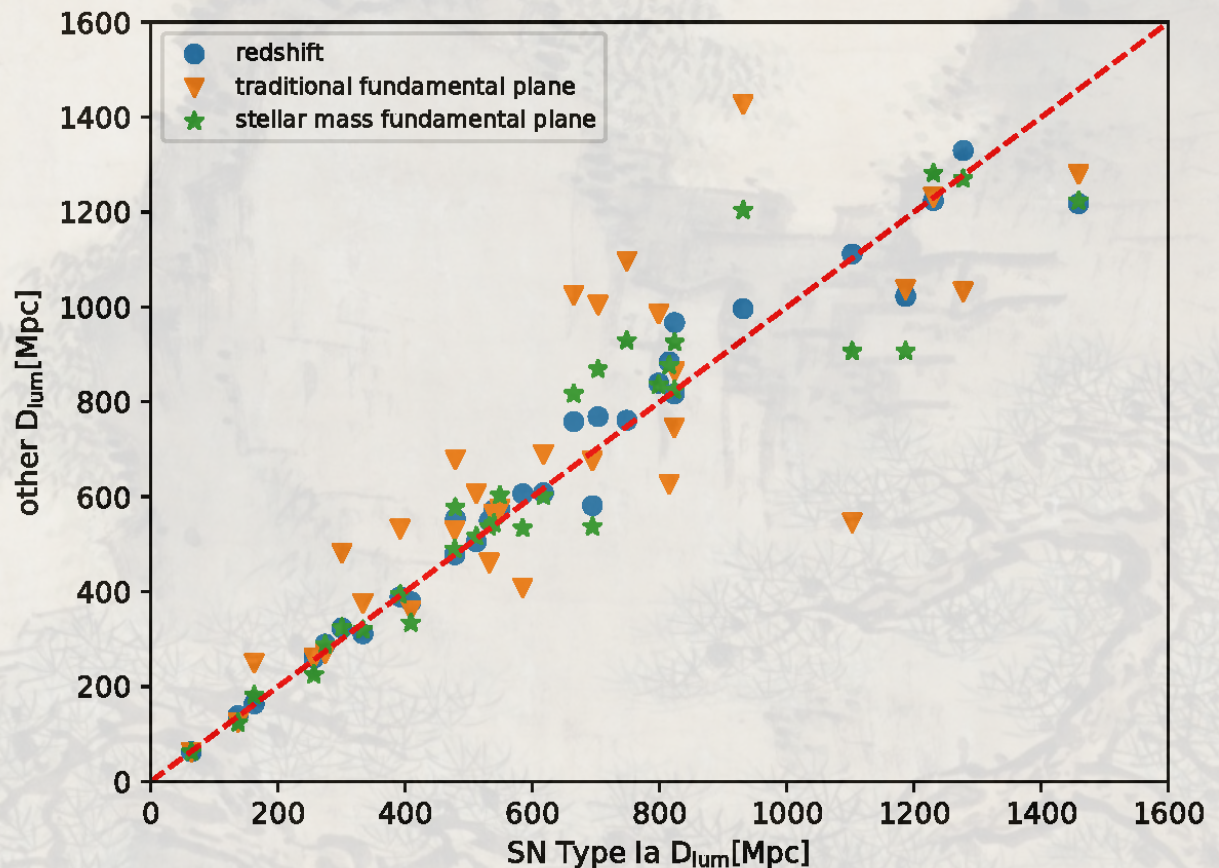
# Comparison to SN Type Ia distances

- Using 740 SN Type Ia distances (Betoule+2014)
- 31 SN are located within our sample of ETG

➤ Scatter Sni Ia vs.

- Redshifts: 8.1%
- Trad. FP: 28.9%
- SM. FP: 14.9%

➤ Results are not as clear as expected, but small sample





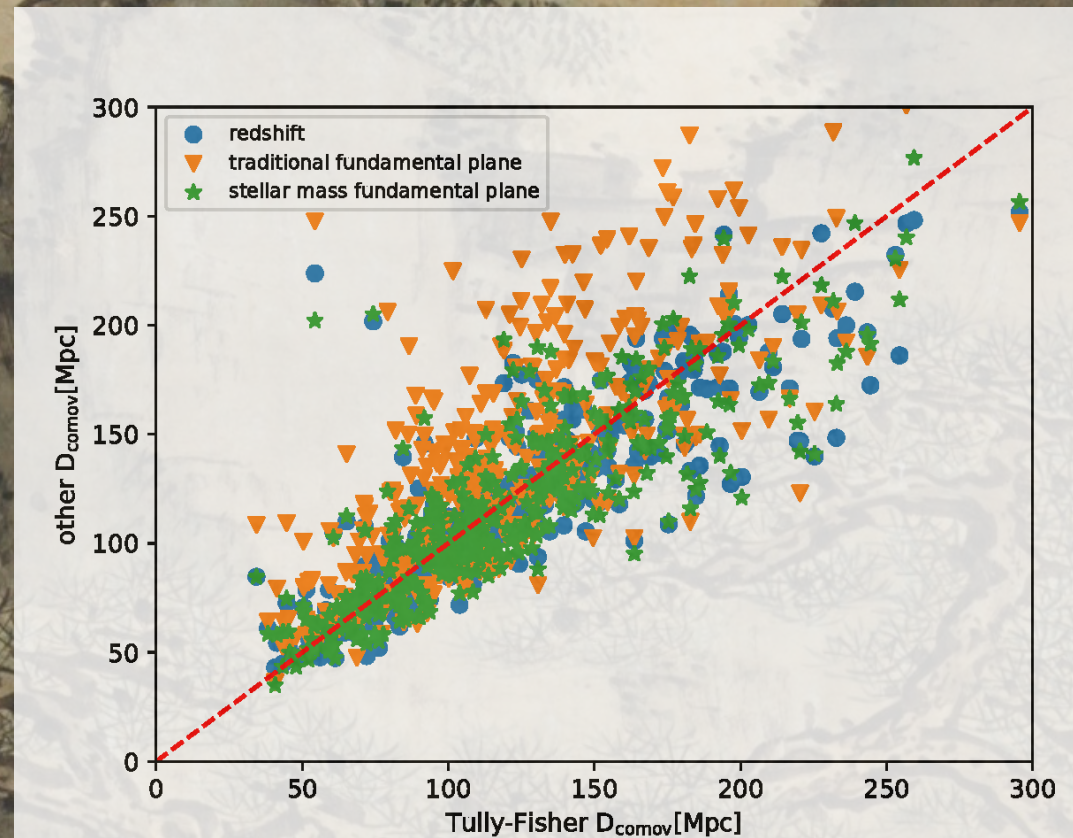
# Comparison to the Tully-Fisher relation distances

- Our group catalogue allows for a comparison of FP distances and TF-relation distances (from NED)
- 397 groups have at least 1 FP and 1 TF galaxy

➤ Scatter Tully-Fisher vs.

- Redshifts: 22.2%
- Trad. FP: 39.5%
- SM. FP: 22.6%

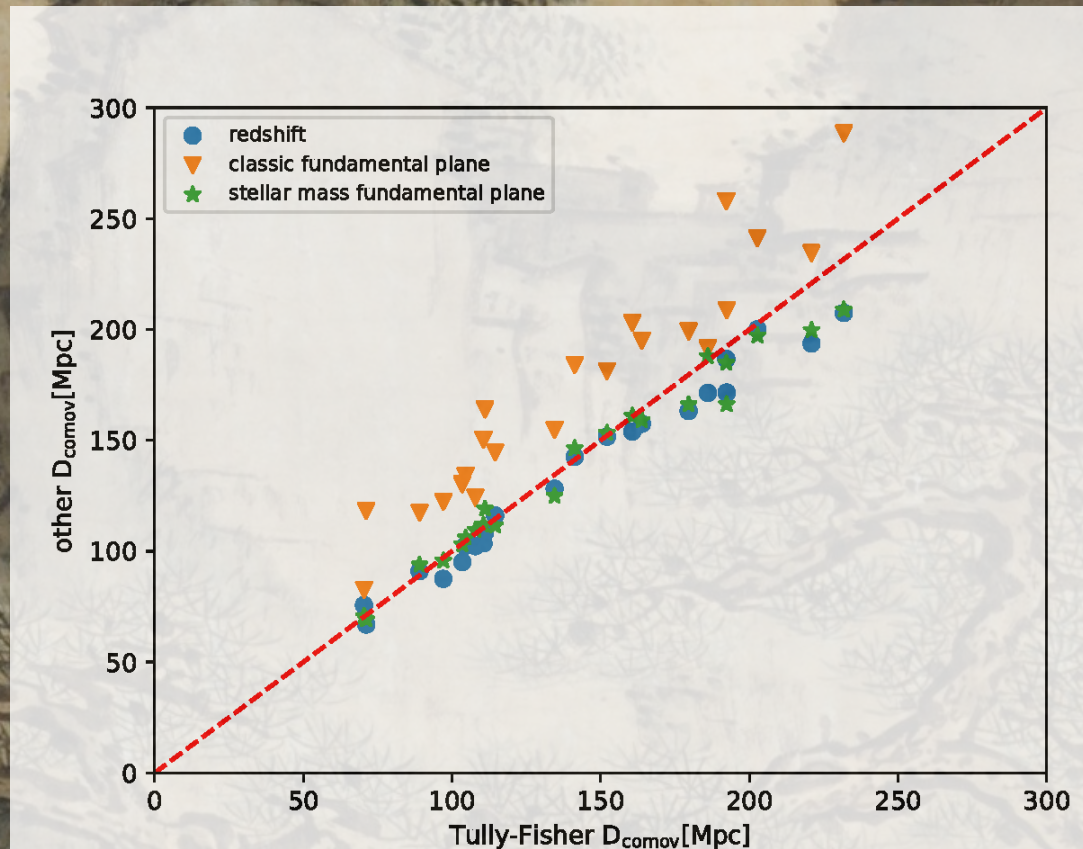
- Systematic bias of Trad FP
- Interlopers may affect these results





# Selecting rich clusters for a better comparison

- Selecting a sub-sample of rich clusters:  
at least 3 FP galaxies and at least 3 TF galaxies
- 22 clusters found
- Scatter Tully-Fisher vs.
  - Redshifts: 6.1%
  - Trad. FP: 27.3%
  - SM. FP: 4.6%
- SM FP agrees better with TF distance than with redshift distances





# Peculiar velocities

- Self-consistent set of (mostly) redshift-independent distances
- Redshifts for the same objects
- → peculiar velocities
- Quality selection (we will focus on clusters, because they provide more solid distances)
- Take care of systematics (difficult)
- Long-term goal: study momentum power spectrum and the  $\beta$  parameter
- Work in progress

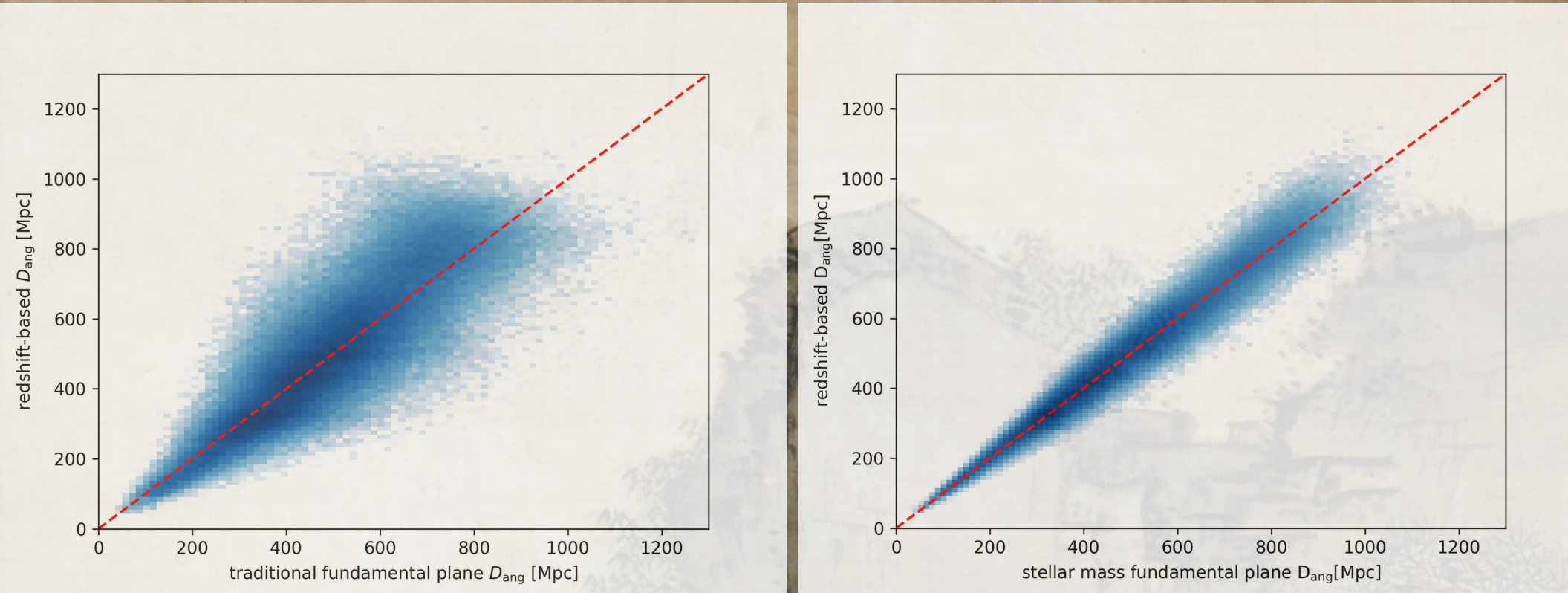


# Summary

- Largest sample of fundamental plane distances (~280 000 galaxies)
- Updated combined SDSS/BOSS/2MRS group catalogue up to  $z=0.5$  (>1 000 000 groups)
- Traditional fundamental plane is biased
- Stellar mass fundamental plane provides notably better distances: 7.7% accuracy
- Good agreement of the SM FP distances with Tully-Fisher distances from NED
- Peculiar motions are work in progress
- **Paper (Saulder+) is about to be submitted**



# ANY QUESTIONS?



감사합니다 !



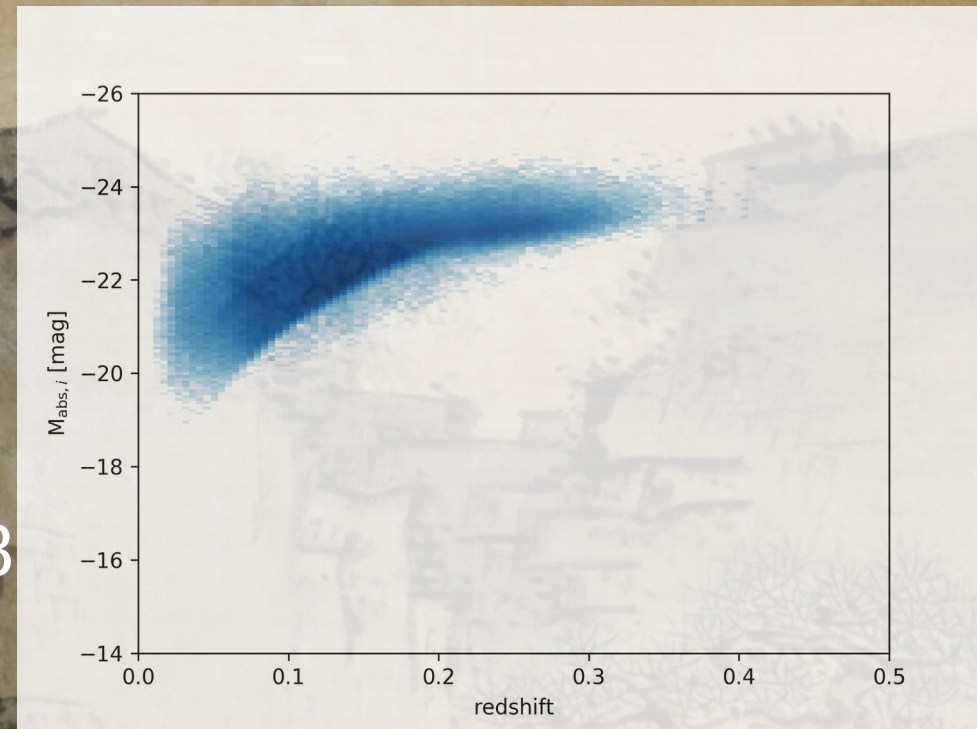
**ADDITIONAL SLIDES  
FOR  
POSSIBLE QUESTIONS**





# Dataset for fundamental plane calibrations

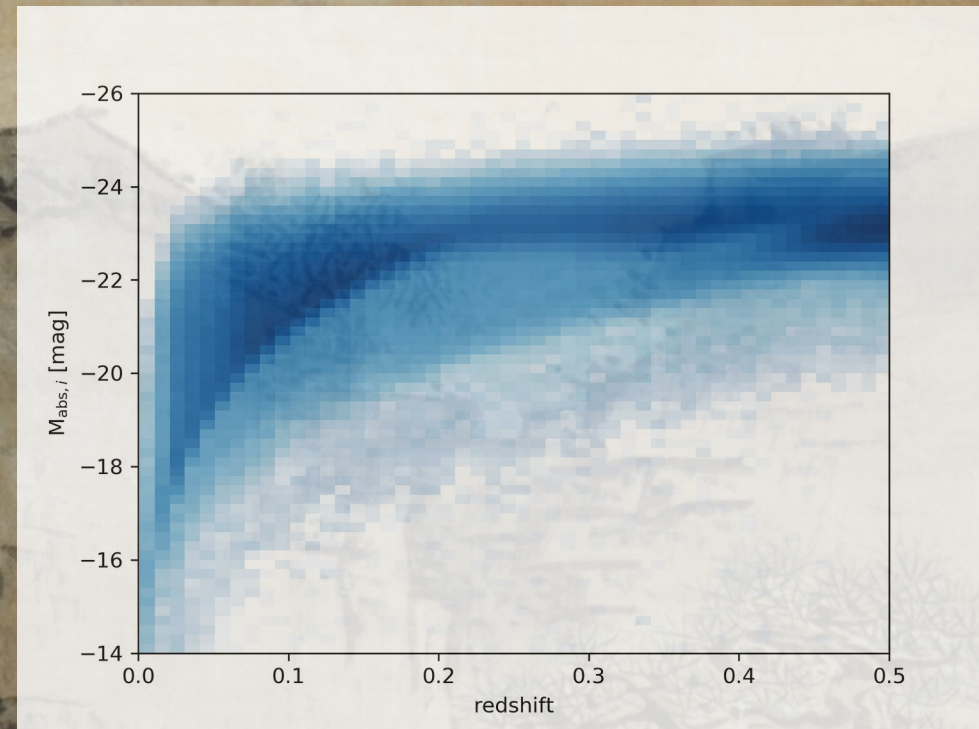
- ~280 000 early-type galaxies from SDSS DR14
- Axis ratio  $< 0.7$
- $(g-r)$  colour  $> 0.65$  mag
- $g > r > i > z$  (the redder, the brighter)
- De Vaucouleur profiles more likely than exponential profile
- De Vaucouleur fitting fraction  $> 0.8$
- Within  $3\text{-}\sigma$  of the red sequence
- Corrected central velocity dispersion  $\in [100 \text{ km/s}, 420 \text{ km/s}]$
- Absolute magnitude  $\in [-19 \text{ mag}, -25.5 \text{ mag}]$
- $\text{Log}(R_0/\text{kpc}) \in [-0.5 \text{ dex}, 1.5 \text{ dex}]$
- Corrected  $(g-r)$  colour  $< 2.5$  mag





# Dataset for group catalogue

- All Galaxies and QSO with spectroscopic redshifts in SDSS DR14
  - SDSS main galaxy sample
  - SDSS LRG sample (low and high  $z$ )
  - BOSS low  $z$  sample
  - CMASS sample



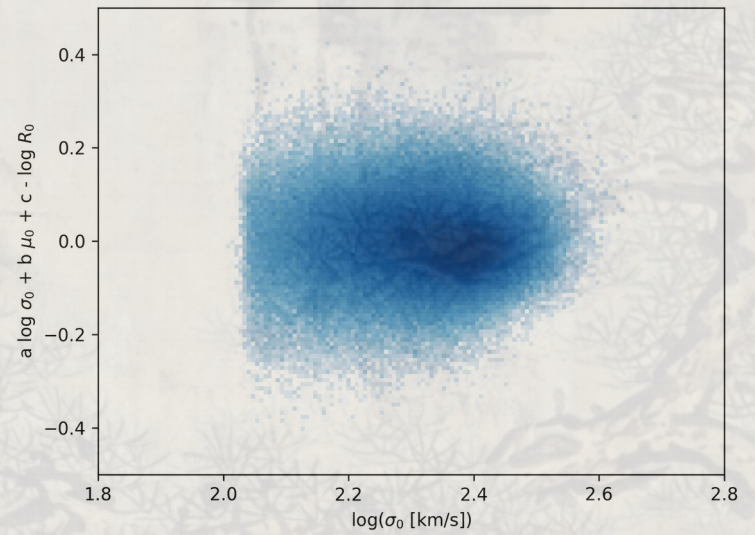
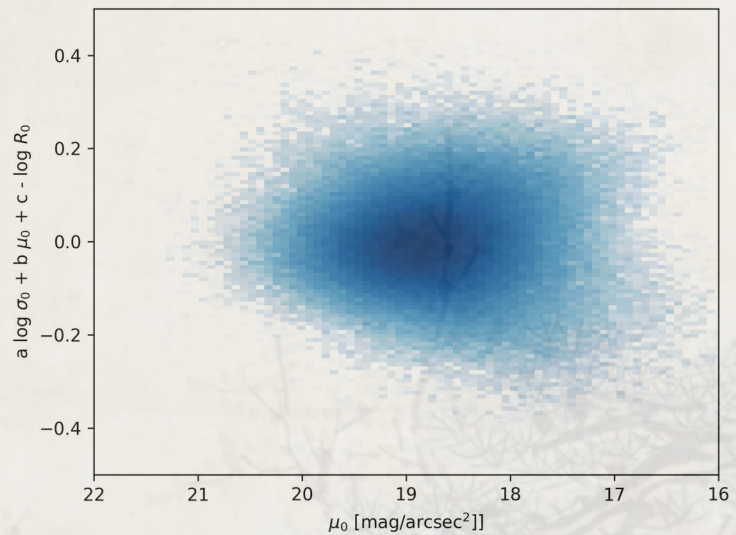
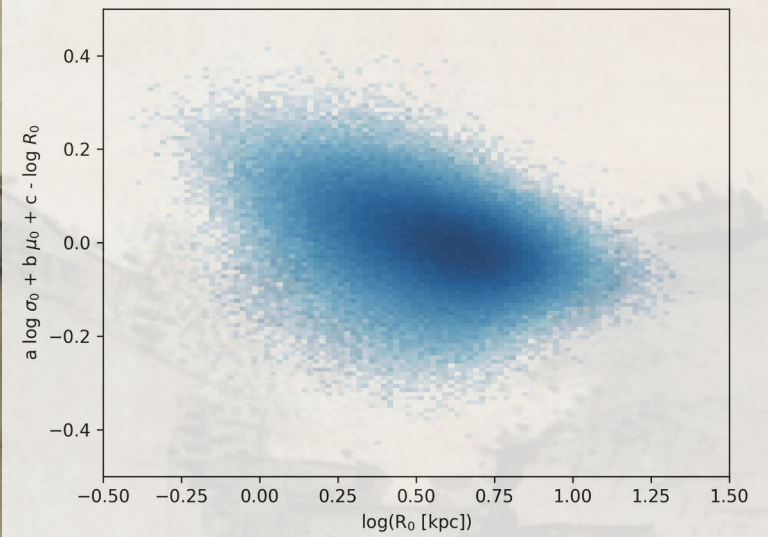
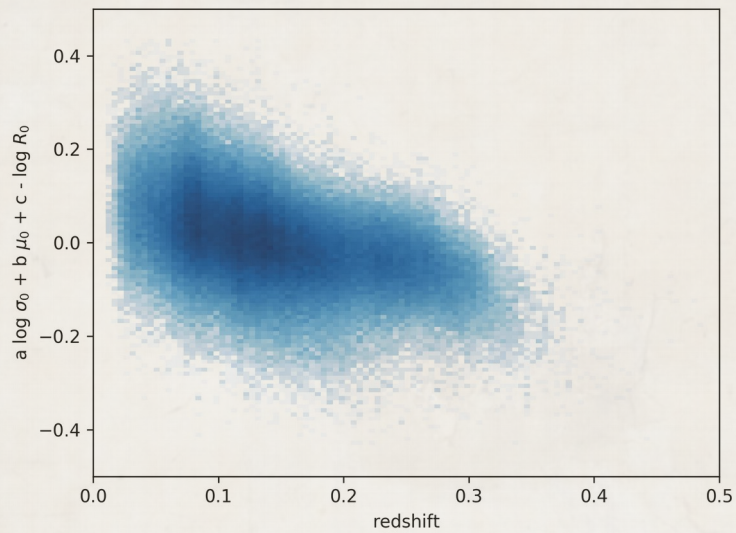
- All galaxies 2MRS within 1 degree of the SDSS DR14 spectroscopic footprint
- 1 269 405 objects as the basis for our group catalogue





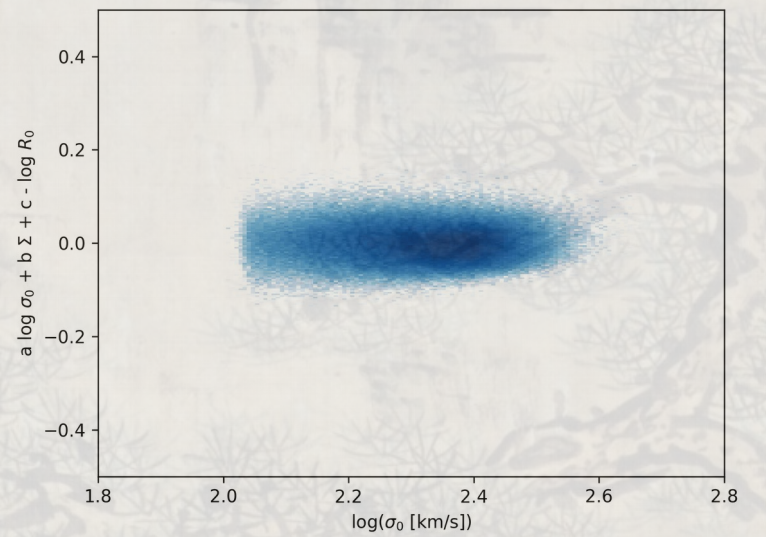
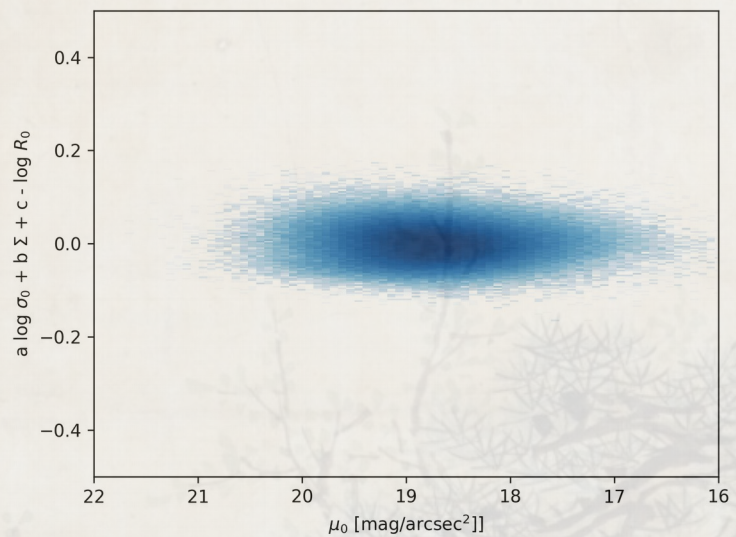
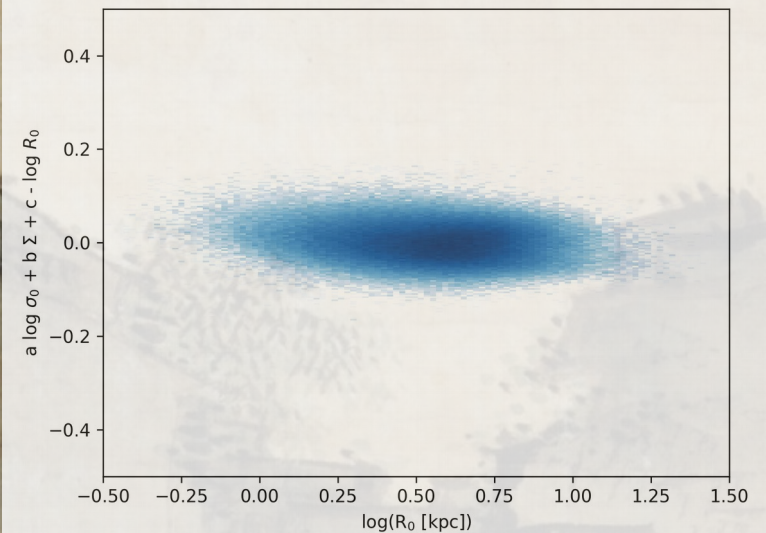
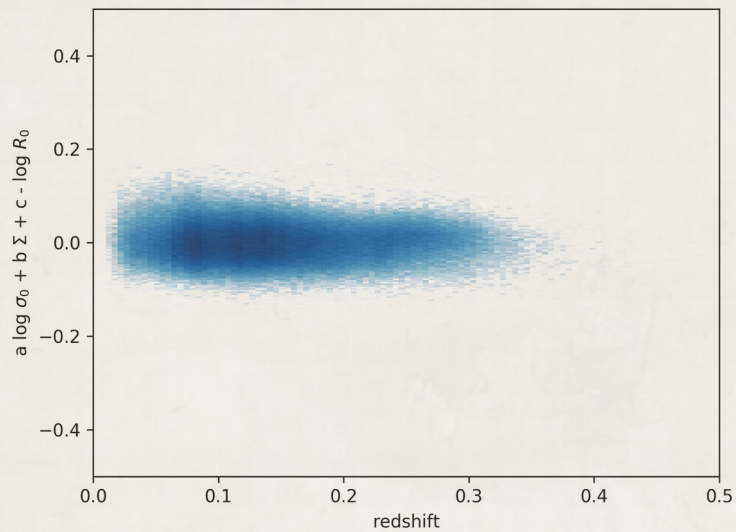


# Residuals of the traditional fundamental plane





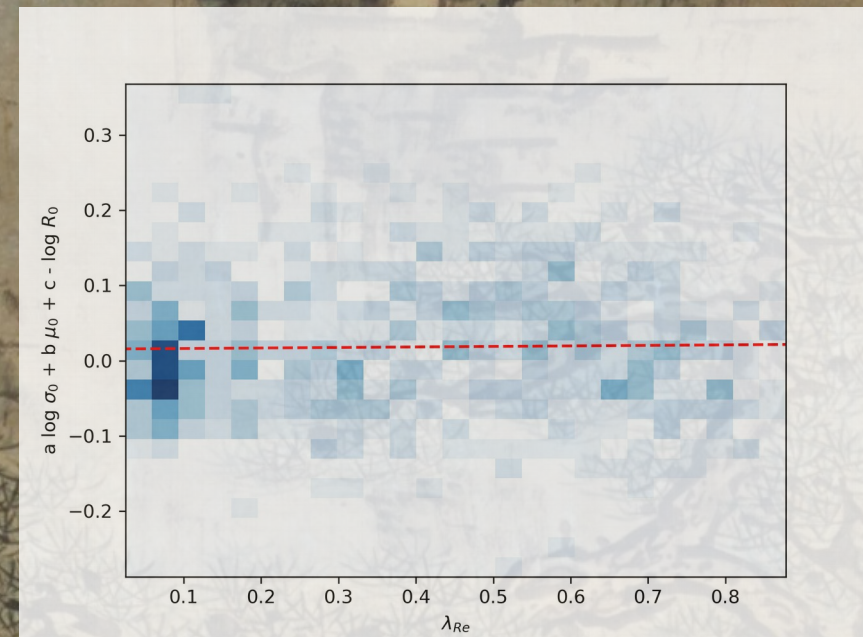
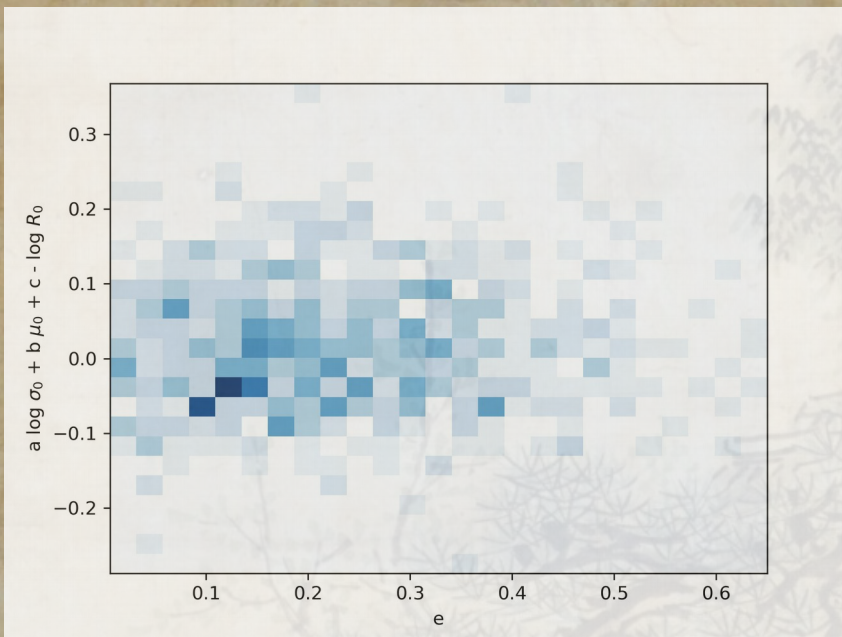
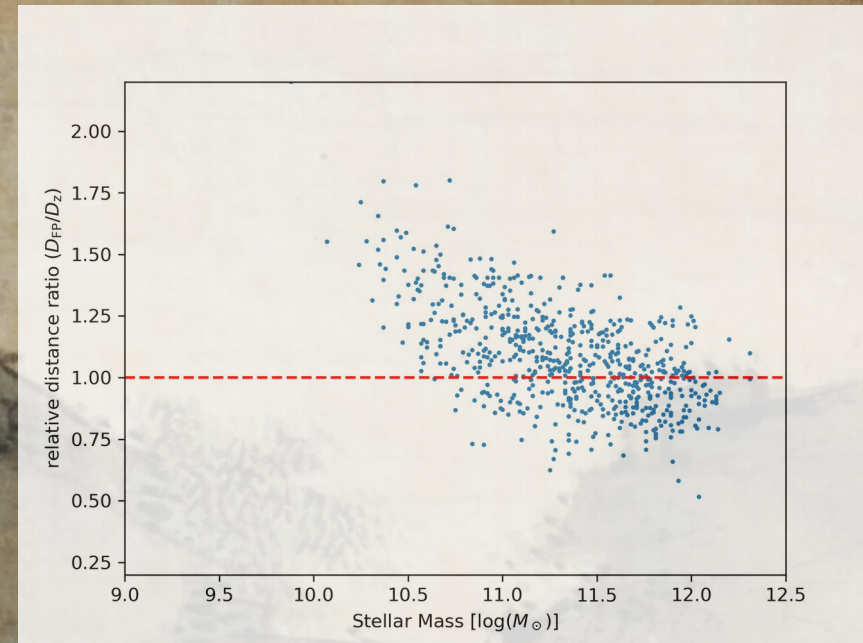
# Residuals of the stellar mass fundamental plane





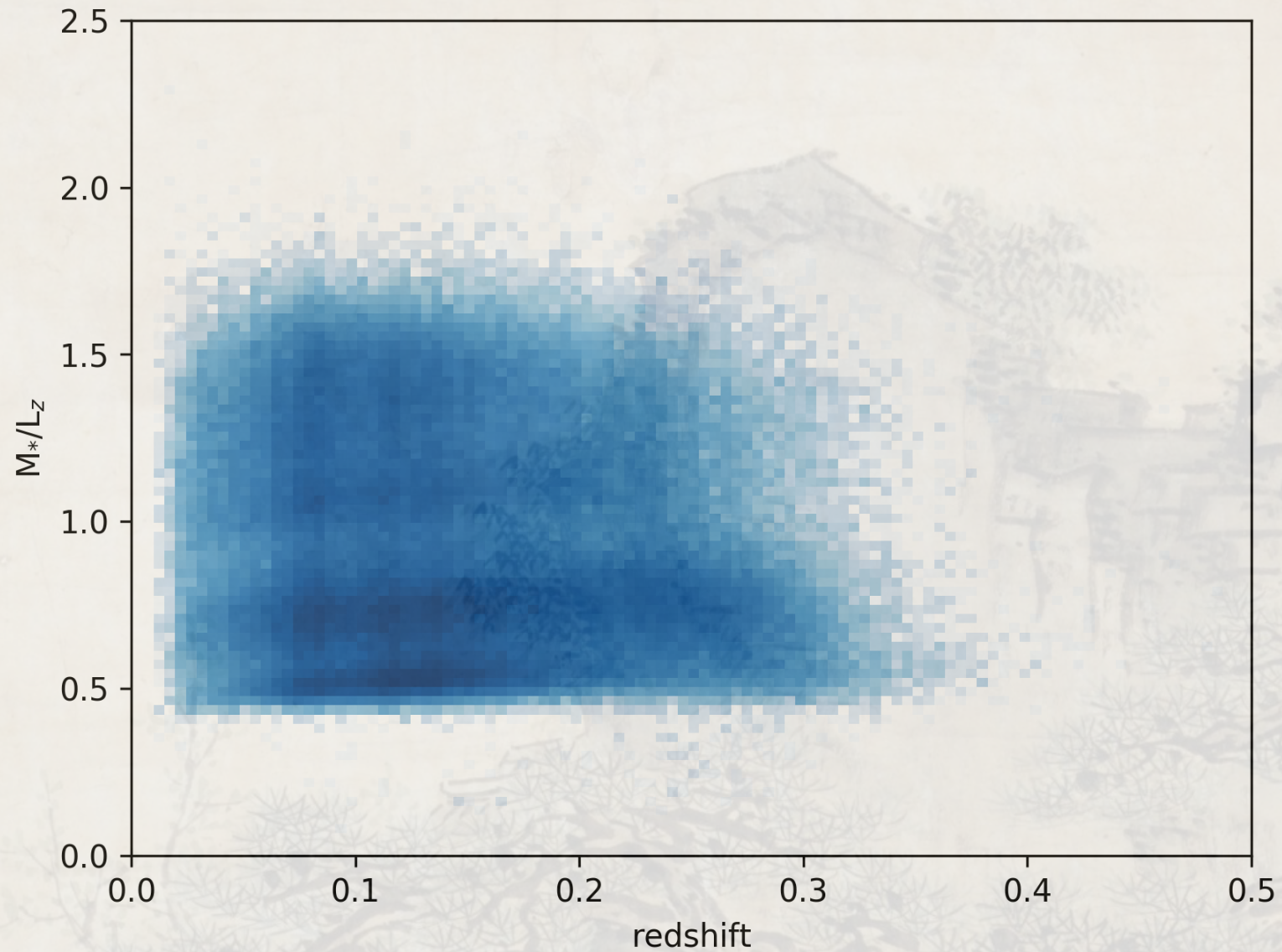
# MaNGA data and residuals

- For the traditional FP
  - Strong trend with stellar mass
  - No correlation with the rotation parameter  $\lambda_{Re}$  or the ellipticity  $e$





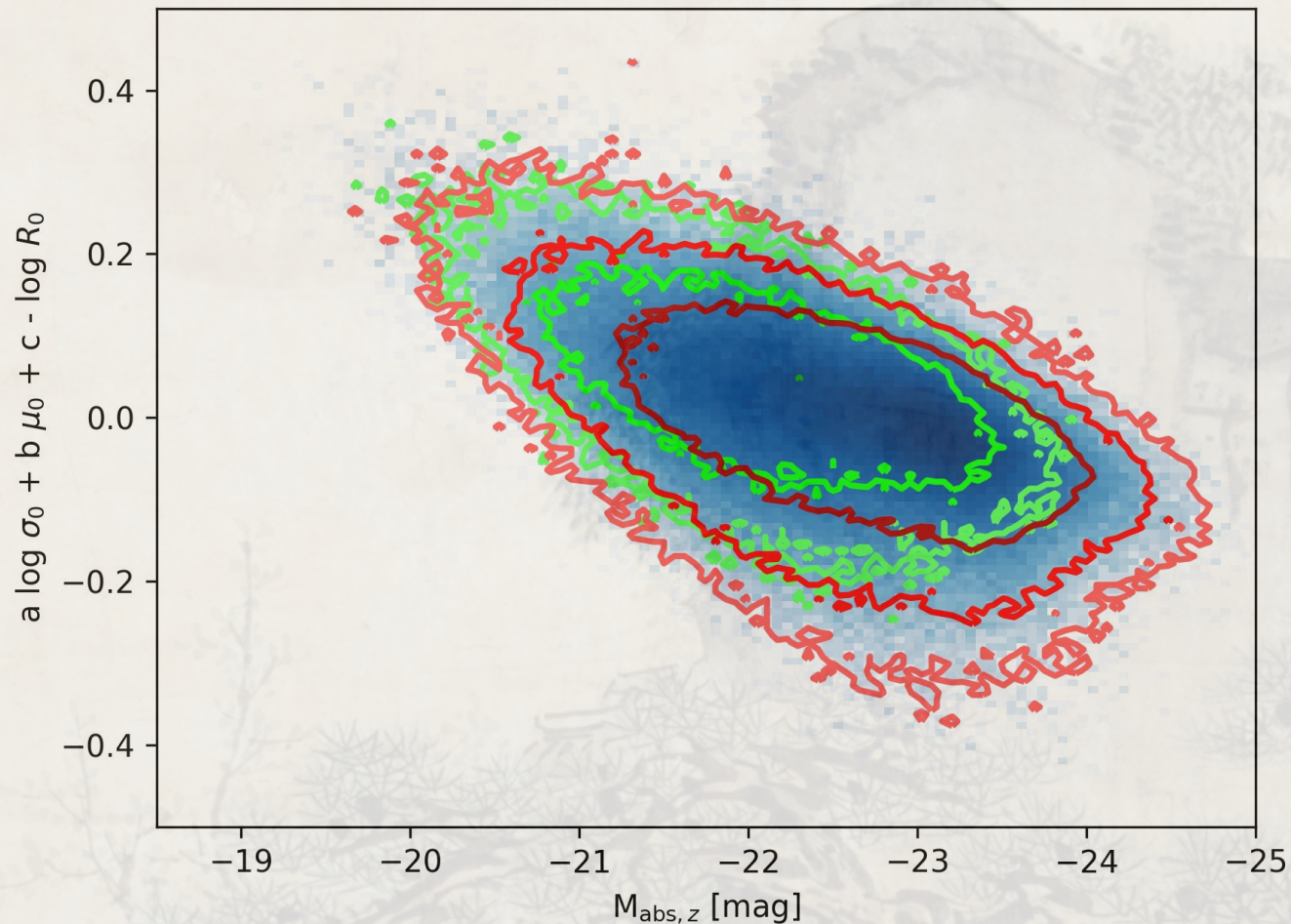
# Stellar mass to light ratio





# Environmental effects

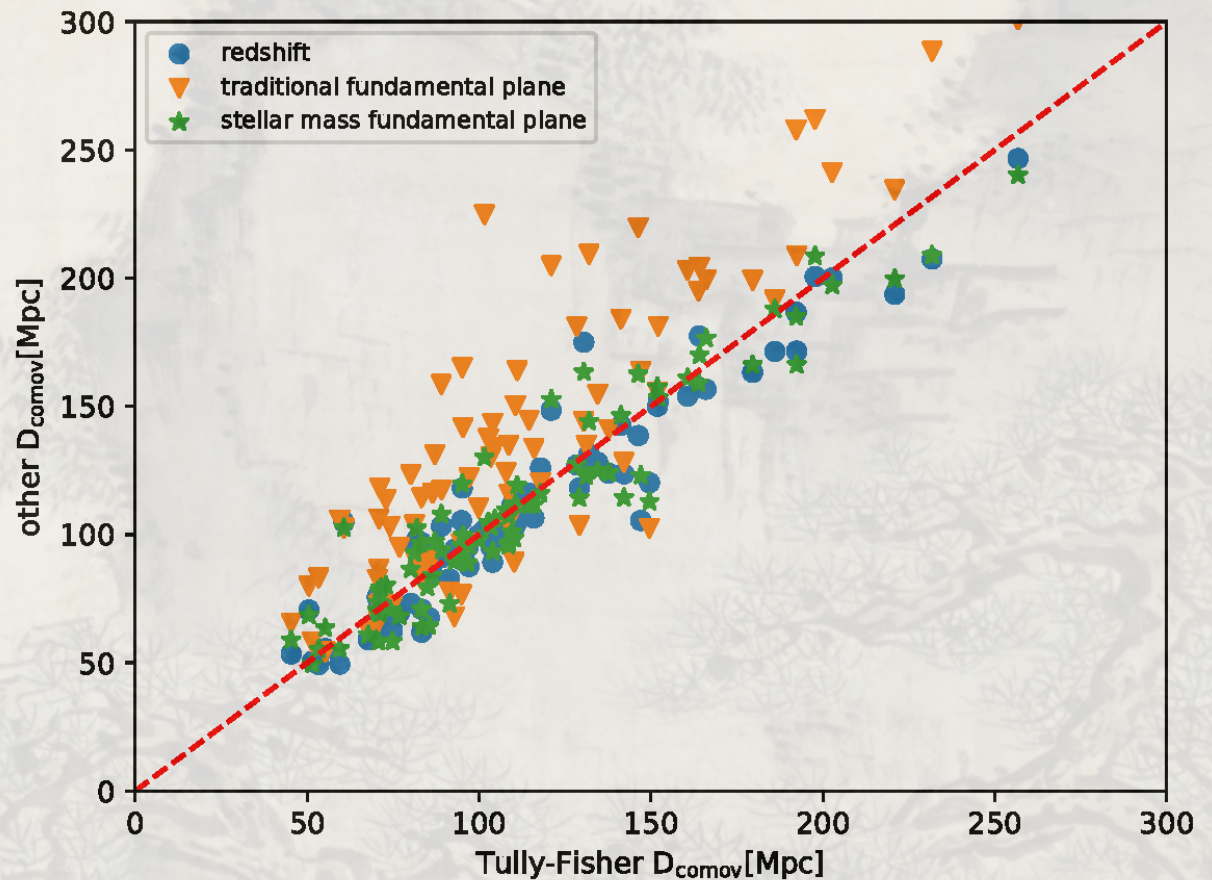
- Centrals are off-set from satellites, but they are also systematically brighter





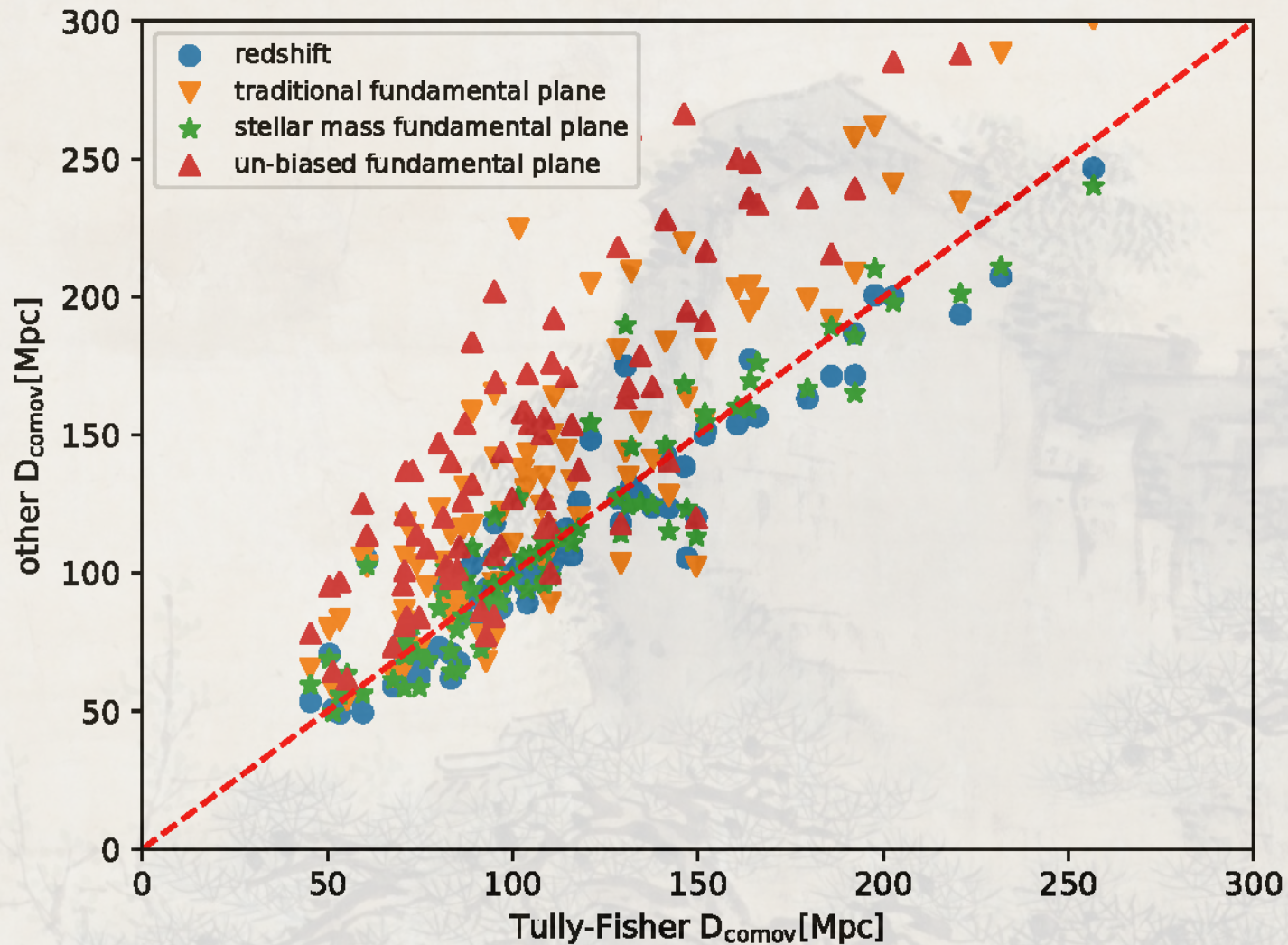
# Tully-Fisher vs. others

- 2 or more TF galaxies and at least 1 FP galaxy
- 84 groups
- Scatter Tully-Fisher vs.
  - Redshifts: 12.0%
  - Trad. FP: 33.3%
  - SM. FP: 13.2%





# Malmquist bias correction vs. Tully-Fisher relation





# The complete picture

- Painting of a gathering of scholars
- Anonymous painter
- Joseon dynasty
- National Museum of Korea
- <https://www.museum.go.kr/site/eng/relic/search/view?relicId=1012>





# N/A

Sorry, but I haven't prepared a slide for this specific question.