## The multivariate halo bias in the eye of a machine

#### Jiaxin Han Shanghai JiaoTong University & Kavli IPMU

Yin Li, Yipeng Jing (SJTU), Takahiro Nishimichi, Wenting Wang and Chunyan Jiang (SHAO)

Arxiv: 1802.0917, MNRAS

#### Introduction

## Halo Bias

• Bias: unfair selection



- halo field is a biased sampling of the density field  $\delta_h = b\delta_m \longrightarrow \rho_h = k\rho_m + (1-b)\overline{\rho}_h$
- depend on all variables that can modify the sampling: multivariate

 $b(M,c,e,j,t_f,\ldots)$ 

- Previous works: marginalized dependences
  - Univariate: well studied b(M) $b(M) = \int b(M, c, e, j, t_f, ...) dP(c, e, j, t_f, ...)$
  - Bivariate: detected "assembly bias"

 $b(M,c), \ b(M,j), \ b(M,t_f)...$ 

Method

## Efficient bias measurement

• Ensemble estimator of bias

$$\delta_h = b\delta_m \; \Rightarrow \; \xi_{hm} = b\xi_{mm}$$

Correlation function: average density profile

$$\xi_{hm} = <\delta_h \delta_m >$$
$$= <\delta_m |h>$$

• individual bias estimator

$$\beta(r) = \delta_m(r) / \xi_{mm}(r)$$

 $\Rightarrow b = <\beta>$ 

## Single-variable dependence $b = < \beta >$



 $M_{\rm vir} > 10^{11.5} M_{\odot} / h \ (500 \ {\rm particles})$ 

Data: LCDM simulation, 3072^3 particles, 600Mpc/h box

### **Bivariate dependence**



### **Bivariate dependence**



# Residual dependence and multivariate estimators



# Residual dependence and multivariate estimators



# Residual dependence and multivariate estimators



## Assembly History Dependence



## MAH dependence

Early-forming haloes: extra dependence on recent MAH



 $a_f$  : the scale factor when the halo is a fraction f of its final mass

## The density profile dependence



## One environment to rule them all



env  $\delta_e$ bias

## Conclusions

- Mass, concentration (formation time), spin, shape are non-redundant bias variables
  - Bias depends mostly on mass at high mass,
  - mostly on (recent) formation history for early-forming halos
  - Lowmass lateforming halos care more about spin and shape.
- $V_{\rm max}/V_{\rm vir}$  is a lossy proxy of formation time, while spin and shape are different
- The environment around halos captures most of the bias dependence on halo structure
  - Defined at as small as 1~2Mpc/h scale

## Summary



arxiv: 1802.0917

### **Bias scale**



## **Gaussian Process Regression**

- Fitting Gaussian random field to maps—constrained Gaussian random field
  - A Gaussian random field: the joint PDF of the values at any field points are multivariate Gaussian
  - Characterized by correlation function
  - Can almost match anything  $\rightarrow$  universal fitting tool

